

$$q = Q - \frac{P^2}{2m^2g}$$

So New hamiltonian is -

$$H'(P, Q) = \frac{P^2}{2m} + mgq$$

$$H'(P, Q) = \frac{P^2}{2m} + mg \left(Q - \frac{P^2}{2m^2g} \right)$$

A-12
Q.11 find generating function $F(P, q)$ for transformation

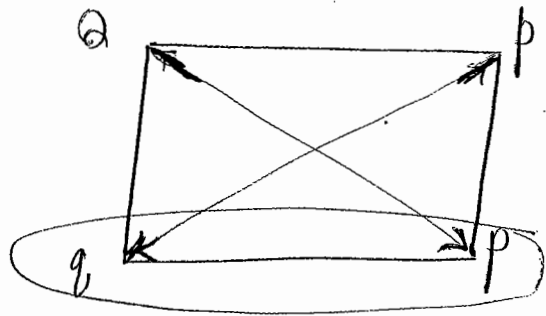
$$p = \frac{1}{Q}, \quad q = PQ^2 \text{ is - ?}$$

- (a) \sqrt{Pq} (b) $-\sqrt{Pq}$ (c) $2\sqrt{Pq}$ (d) $-2\sqrt{Pq}$.

Solⁿ Given transformation relation is -

$$p = \frac{1}{Q}$$

$$q = PQ^2$$



$$p = \frac{\partial F(q, P)}{\partial q} \quad \text{--- (I)}$$

$$Q = \frac{\partial F(Q, p)}{\partial p} \quad \text{--- (II)}$$

from (I) relation -

$$\sqrt{\frac{p}{q}} = \frac{\partial F}{\partial q}$$

$$\sqrt{p} = 2\sqrt{q} + c_1(p) = F \quad \text{--- (a)}$$

from (II)

$$\sqrt{\frac{q}{p}} = \frac{\partial F}{\partial p} \Rightarrow F = \sqrt{q} \cdot 2\sqrt{p} + c_2(q) \quad \text{--- (b)}$$

So from (a) & (b) $|F = 2\sqrt{Pq}|$ R

Q9 Values of 'a' and 'b' for which following transformations

$Q = (2q)^a \cos^b p$, $P = (2q)^a \sin^b p$ is canonical are -

- (A) $a = \frac{1}{2}, b = 1$ (B) $a = 2, b = \frac{1}{2}$ (C) $a = 1, b = 1$, (D) $a = \frac{1}{2}, b = \frac{1}{2}$

Solⁿ For canonical position bracket should be zero.

$$\text{So } \{Q, P\}_{q,p} = 1$$

$$\Rightarrow \{ (2q)^a \cos^b p, (2q)^a \sin^b p \}_{q,p} = 1$$

$$\Rightarrow 2a (2q)^{a-1} \cos^b p \cdot (2q)^a b \sin^{b-1} p \cos p + (2q)^a b \cos^b p \sin p \times 2a (2q)^a \sin^b p = 1$$

$$\Rightarrow 2ab (2q)^{2a-1} \cos^{b-1} p \sin^{b-1} p [\cos^2 p + \sin^2 p] = 1$$

$$2ab (2q)^{2a-1} (\cos p \sin p)^{b-1} = 1$$

\therefore R.H.S is co-ordinate independent \therefore L.H.S.

should also be co-ordinate independent.

So $2a-1 = 0$ and $b-1 = 0$

$$\boxed{a = \frac{1}{2}}$$

$$\boxed{b = 1}$$

* Hamilton Jacobi theory:-

In this theory a

canonical transformation is done in such a way that new hamiltonian becomes zero. And G.F. for this C.T. is chosen to be $F[q, P]$

$$q, p \xrightarrow{\text{C.T.}} Q, P$$

$$H \xrightarrow{F(q, P)} H' = 0$$

Hamilton's equation in new co-ordinate-

$$\dot{Q} = \frac{\partial H'}{\partial P} = 0 \quad \because H' = 0$$

$$\dot{P} = -\frac{\partial H'}{\partial Q} = 0$$

$$\Rightarrow Q = \text{constant}$$

$$\leftarrow P = \text{constant}$$

Importance of $F(q, P, t)$ in H.J. theory.

Total time derivative of F .

$$\frac{dF}{dt} = \frac{\partial F}{\partial q} \dot{q} + \frac{\partial F}{\partial p} \dot{p} + \frac{\partial F}{\partial t}$$

$$\frac{dF}{dt} = \frac{\partial F}{\partial q} \dot{q} + \frac{\partial F}{\partial t}$$

Relation for $f(q, P, t)$

$$p = \frac{\partial F}{\partial q}$$

$$Q = -\frac{\partial F}{\partial P}$$

$$H' = H + \frac{\partial F}{\partial t}$$

$$\frac{\partial F}{\partial t} = -H$$

$$\frac{dF}{dt} = p\dot{q} - H$$

$$\boxed{\frac{dF}{dt} = L} \quad \text{Lagrange's eqn}$$

$$\boxed{F = \int L dt} = S$$

Action (S)

In H-J theory F is equal to action (S).

Imp.
* H-J Equation :-

$$\boxed{H + \frac{\partial S}{\partial t} = 0}$$

In Hamilton every place of p we will write -

$$\boxed{p = \frac{\partial S}{\partial q}}$$

Q. Write hamilton-Jacobi eqⁿ for simple pendulum.

(r, θ)

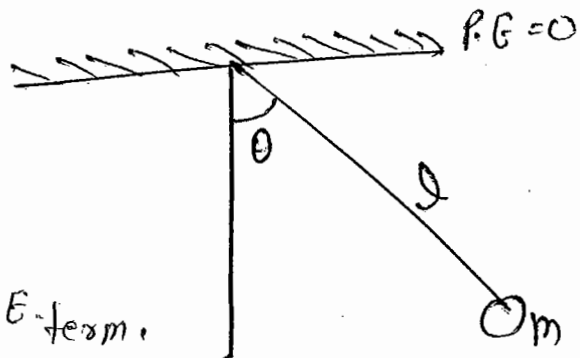
$$r = l$$

$$\dot{r} = 0$$

$$p_r = 0$$

$$H = \frac{p_r^2}{2m} + \frac{p_\theta^2}{2mr^2} + \text{P.E. term.}$$

$$= \frac{p_\theta^2}{2ml^2} - mgl \cos \theta$$



$$p_\theta = \left(\frac{\partial S}{\partial \theta} \right)$$

So H.J. eqⁿ for simple pendulum

$$H + \frac{\partial S}{\partial t} = 0$$

$$\left[\frac{1}{2ml^2} \left(\frac{\partial S}{\partial \theta} \right)^2 - mgl \cos \theta + \frac{\partial S}{\partial t} = 0 \right]^*$$

Non-linear P.D.E.

Solution of Non-linear partial diff. eqⁿ can be found by variable separation on summation not multiplication

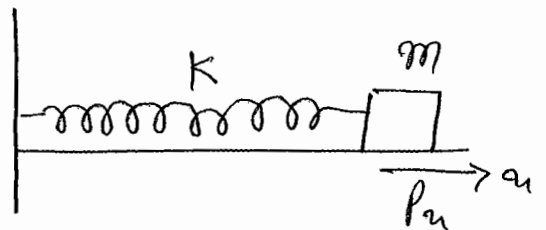
* H-J Equation for S.H.O. :-

$$H + \frac{\partial S}{\partial t} = 0$$

$$p = \frac{\partial S}{\partial q}$$

$$H = \frac{p_u^2}{2m} + \frac{1}{2} k u^2$$

$$= \frac{1}{2} m \left(\frac{\partial S}{\partial u} \right)^2 + \frac{1}{2} k u^2$$



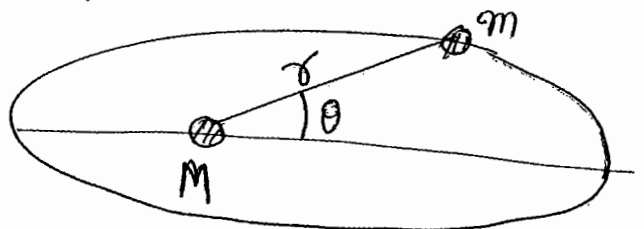
Here H-J eqⁿ -

$$\left[\frac{1}{2} m \left(\frac{\partial S}{\partial u} \right)^2 + \frac{1}{2} k u^2 + \frac{\partial S}{\partial t} = 0 \right]$$

Ques: H-J equation for planetary motion.

Solⁿ :-

$$H = \frac{p_r^2}{2m} + \frac{p_\theta^2}{2mr^2} - \frac{GMm}{r}$$



$$p_r = \left(\frac{\partial S}{\partial r} \right), \quad p_\theta = \frac{\partial S}{\partial \theta}$$

$$\frac{1}{2m} \left(\frac{\partial S}{\partial r} \right)^2 + \frac{1}{2mr^2} \left(\frac{\partial S}{\partial \theta} \right)^2 - \frac{GMm}{r} + \frac{\partial S}{\partial t} = 0$$

↑ Non-Linear P.D.E.

$$S = S(r, \theta, t)$$

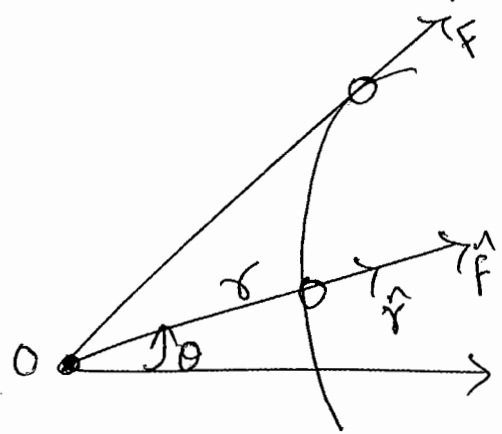
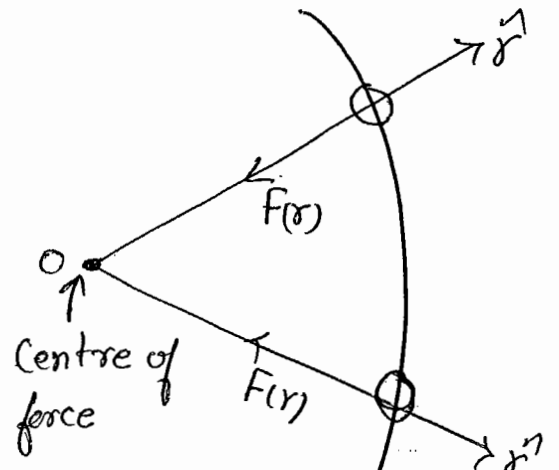
$$S = S_1(r) + S_2(\theta) + S_3(t).$$

Central force Motion :-

force is directed towards or away from a point is central force.

$$\vec{F} = \pm F(r) \hat{r}$$

- When \vec{F} and \hat{r} is in same direction then we use +ve sign.
- When \vec{F} and \hat{r} is in opposite direction then we use -ve sign.



* Properties :-

$$\nabla \times \vec{F} = 0$$

⇒ Central force is conservative.

⇒ Total energy is conserved.

$$\vec{L}_0 = 0, \quad \vec{L}_0 \text{ is conserved.}$$

When angular momentum is conserved then it must be two directional problem.

\vec{F} is conservative. So we can define P.E. ($V(r)$)

$$V(r) = - \int f(r) dr$$

$f(r)$ is +ve for repulsive force
 $f(r)$ is -ve for attractive force

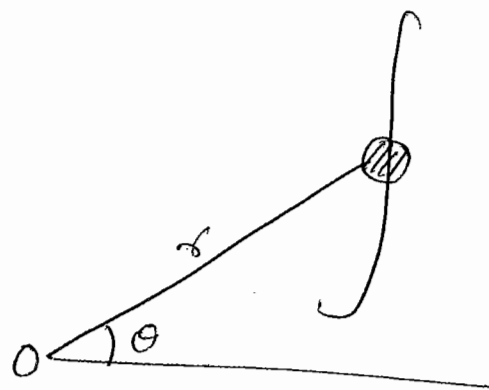
$$\vec{f} = - \nabla V(r)$$

$$f = - \frac{\partial V(r)}{\partial r}$$

Equation of motion under central force :-

$$L = T - V$$

$$L = \frac{1}{2} m (\dot{r}^2 + r^2 \dot{\theta}^2) - V(r)$$



Here L is independent of t so total energy is conserved and L is independent of θ so L_{θ} (angular momentum) is conserved.

Equation of motion :-

r -equation :-

$$\frac{d}{dt} \left(\frac{\partial L}{\partial \dot{r}} \right) - \frac{\partial L}{\partial r} = 0$$

$$m\ddot{r} - m r \dot{\theta}^2 + \frac{\partial V}{\partial r} = 0$$

$$\boxed{m\ddot{r} - m r \dot{\theta}^2 - f(r) = 0} \quad \leftarrow \text{imp. radial eq}^n \text{ of motion.}$$

θ -equation :-

$$\frac{d}{dt} \left(\frac{\partial L}{\partial \dot{\theta}} \right) - \frac{\partial L}{\partial \theta} = 0$$

$$\frac{d}{dt} (m r^2 \dot{\theta}) - 0 = 0$$

$$\frac{d}{dt} (m r^2 \dot{\theta}) = 0$$

$$\boxed{m r^2 \dot{\theta} = \text{Constant}} \quad \text{imp.}$$

$$\boxed{2\dot{r}\dot{\theta} + r\ddot{\theta} = 0} \quad \text{imp.}$$

$$p_{\theta} = \frac{\partial L}{\partial \dot{\theta}} = \boxed{mr^2 \dot{\theta} = l} \text{ Angular momentum.}$$

Q. Transverse velocity of a particle under central force varies with r as -

$$\dot{\theta} = \frac{l}{mr^2}$$

- (a) $\frac{1}{r}$ (b) $\frac{1}{r^2}$ (c) r (d) $\frac{1}{r^2}$

Solⁿ

Transverse velocity = $r\dot{\theta}$

$$= r \cdot \frac{l}{mr^2} = \frac{l}{m} \cdot \frac{1}{r}$$

$\left. \begin{aligned} \dot{\theta} &= \frac{l}{mr^2} \\ \text{k.E.} &= \frac{1}{2} m \left(\underbrace{\dot{r}^2}_{\text{Radial Velocity}} + \underbrace{r^2 \dot{\theta}^2}_{\text{Trans. velocity}} \right) \end{aligned} \right\}$

Hence transverse velocity $\propto \frac{1}{r}$

* Equation of path in central force motion :-

is independent of ~~path~~ time.

remove t from equation of motion.

Introduce $u = \frac{1}{r}$

This leads to following eqⁿ

$$\boxed{\frac{d^2 u}{d\theta^2} + u = -\frac{mf}{u^2 u^2}}$$

Diff. eqⁿ of path

Use! - It is use to find $f(r)$ when relation b/w r and θ is given.

$$\begin{aligned} u &= \frac{1}{r} \Rightarrow r = \frac{1}{u} \\ \dot{r} &= -\frac{1}{u^2} \frac{du}{dt} \\ &= -\frac{1}{u^2} \frac{du}{d\theta} \cdot \frac{d\theta}{dt} \\ &= -r^2 \frac{du}{d\theta} \cdot \dot{\theta} \\ &= -r^2 \frac{du}{d\theta} \cdot \frac{l}{mr^2} \\ &= -l \frac{du}{d\theta} \end{aligned}$$

Q. If $r = a \cos \theta$ for a particle moving under central force what is force law?

Solⁿ

$$u = \frac{1}{r}$$

$$u = \frac{1}{a} \sec \theta$$

$$\frac{du}{d\theta} = \frac{1}{a} \sec \theta \tan \theta$$

$$\frac{d^2u}{d\theta^2} = \frac{1}{a} [\sec \theta \tan^2 \theta + \sec^3 \theta]$$

Differential equation of path -

$$\frac{d^2u}{d\theta^2} + u = \frac{-mf}{l^2 u^2}$$

$$\frac{u \tan^2 \theta + u \sec^2 \theta + u}{1} = \frac{-mf}{l^2 u^2}$$

$$2u \sec^2 \theta = \frac{-mf}{l^2 u^2}$$

$$f = \frac{-2l^2 u^3 \sec^2 \theta}{m}$$

$$f = \frac{-2l^2 a^2}{m} u^5$$

$$f = \frac{-2l^2 a^2}{m} \frac{1}{r^5}$$

$$\boxed{f \propto \frac{1}{r^5}} \quad \leftarrow \text{Force Law}$$

Note ①

If $r = a \sin \theta$

$$\Rightarrow \boxed{f \propto \frac{1}{r^5}}$$

②

if $r^n = a \cos n\theta$
or $r^n = a \sin n\theta$

$$\boxed{f \propto \frac{1}{r^{2n+3}}}$$

Imp.
Q. A particle is moving in circle of radius R under a force which is always directed towards a point on periphery of circle.
 (i) What is the force law. (ii) What is total energy.

Soln

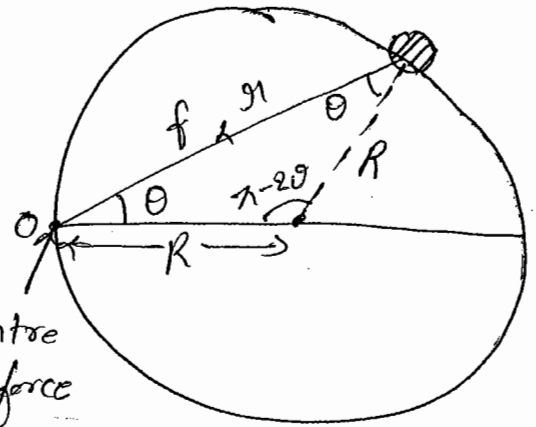
By sine law

$$\frac{r}{\sin(\pi - 2\theta)} = \frac{R}{\sin\theta}$$

$$\frac{r}{\sin 2\theta} = \frac{R}{\sin\theta} \Rightarrow \frac{r}{2\sin\theta\cos\theta} = \frac{R}{\sin\theta}$$

Centre of force

$$\boxed{r = 2R\cos\theta}$$



Q. A particle is moving under a central force if eqn of path of the particle is $r = Ae^{k\theta}$. What is potential Energy of the particle.

Soln

$$r = Ae^{k\theta}$$

is eqn of path becoz this relation b/w two co-ordinate and also it is time independent. when it is written in differential form then it is klq differential eqn of path.

Ans: $f \propto \frac{1}{r^3}$

$$V(r) = \frac{-(k^2 + 1)l^2}{2mr^2}$$

T.E. = 0

for total energy:

T.E. = K.E. + P.E.

$$= \frac{1}{2} m (\dot{r}^2 + r^2 \dot{\theta}^2) + P.E. \quad \left| \begin{array}{l} r = Ae^{k\theta} \\ \dot{r} = kAe^{k\theta} \dot{\theta} \\ \dot{r} = kr\dot{\theta} \end{array} \right. \quad \dot{\theta} = \frac{l}{mr^2}$$



* Effective Potential :- [Potential Energy] :-

It is introduced to convert 2-D problem (in central force motion) into 1-D problem.

Total Energy -

$$E = K.E + P.E. \Rightarrow E = \frac{1}{2}mv^2 + V(r)$$

$$= \frac{1}{2}m(\dot{r}^2 + r^2\dot{\theta}^2) + \underbrace{V(r)}_{\substack{\uparrow \\ \text{Actual P.E.}}}$$

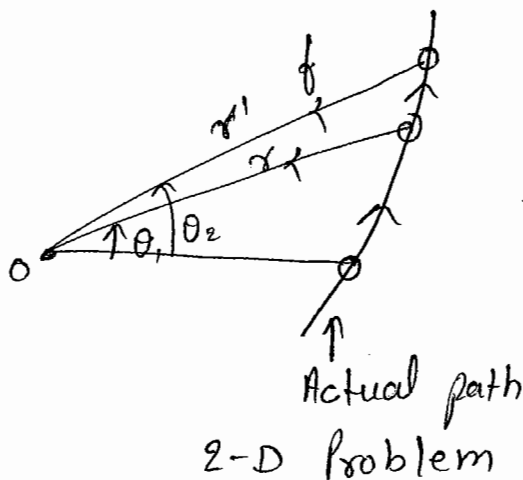
Put $\dot{\theta} = \frac{l}{mr^2}$

$$E = \underbrace{\frac{1}{2}m\dot{r}^2}_{K.E.} + \underbrace{\frac{l^2}{2mr^2} + V(r)}_{P.E. \text{ or Eff. Pot.}}$$

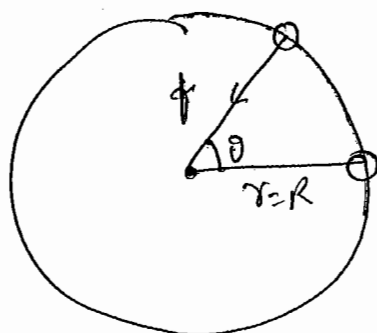
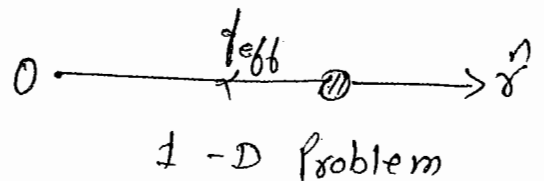
$$E = \frac{1}{2}m\dot{r}^2 + V_{\text{eff}}(r)$$

$$V_{\text{eff}} = V(r) + \frac{l^2}{2mr^2}$$

$$f_{\text{eff}} = -\frac{\partial V_{\text{eff}}}{\partial r}$$

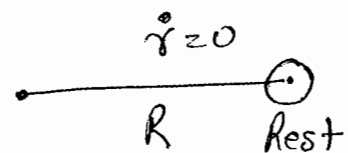


\Rightarrow



2-D circular motion

\Rightarrow



$$f_{\text{eff}} = 0 \quad | \quad E = V_{\text{eff}}$$

Q.34

Solⁿ It is circular motion

$$\therefore V_{\text{eff}} = 0$$

$$\frac{\partial V_{\text{eff}}}{\partial r} = 0$$

$$\frac{\partial}{\partial r} \left(\frac{-k}{r} + \frac{l^2}{2mr^2} \right)_{r=r_0} = 0$$

$$\left[\frac{+k}{r^2} + \left(-\frac{l^2}{mr^3} \right) \right]_{r=r_0} = 0$$

$$\frac{k}{r_0^2} = \frac{l^2}{mr_0^3}$$

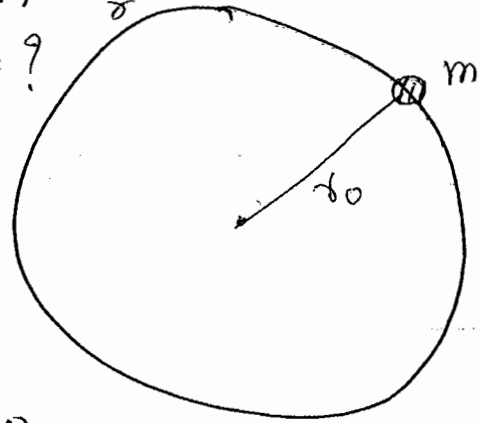
$$l^2 = kmr_0$$

$$\boxed{l = \sqrt{kmr_0}} \quad \underline{\underline{A_2}}$$

B.A.5
Q.20

$$V(r) = \frac{-k}{r}$$

$$l = ?$$

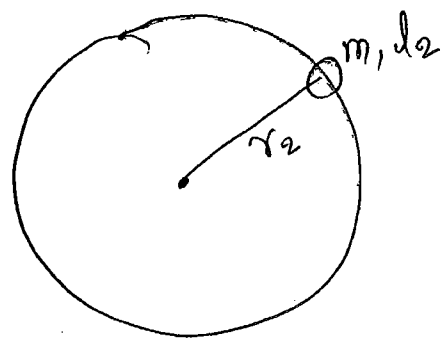
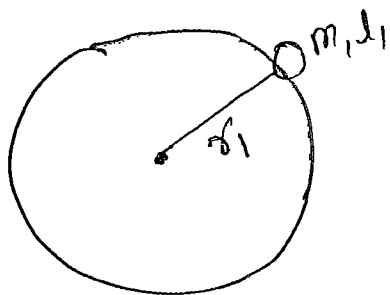


Solⁿ

$$\text{Given } \Rightarrow \frac{l_1}{l_2} = 2$$

$$\frac{r_1}{r_2} = ?$$

$$V(r) = \frac{1}{2} kr^2$$



∴ motion is circular motion.

$$\text{So } V_{\text{eff}} = 0$$

$$\frac{\partial V_{\text{eff}}}{\partial r} = 0$$

$$\frac{\partial}{\partial r} \left(\frac{1}{2} Kr^2 + \frac{l^2}{2mr^2} \right) = 0$$

$$Kr - \frac{l^2}{mr^3} = 0$$

$$l = \sqrt{Kmr^4}$$

$$l = \sqrt{Km} r^2$$

$$l_1 = \sqrt{Km} r_1^2 \quad \text{--- (i)}$$

$$l_2 = \sqrt{Km} r_2^2 \quad \text{--- (ii)}$$

$$\frac{l_1}{l_2} = \frac{\sqrt{Km} r_1^2}{\sqrt{Km} r_2^2}$$

$$2 = \frac{r_1^2}{r_2^2}$$

$$\text{So } \boxed{\frac{r_1}{r_2} = \sqrt{2}} \quad \underline{\underline{\text{Ans}}}$$

Q.3

Solⁿ

$$V(r) = Kr^3$$

$$E = \text{K.E.} + \text{P.E.} \quad \text{--- (i)}$$

But in circular motion -

$$E = V_{\text{eff.}} \quad \text{--- (ii)}$$

$$\text{K.E.} + \text{P.E.} = \frac{l^2}{2mr^2} + V(r)$$

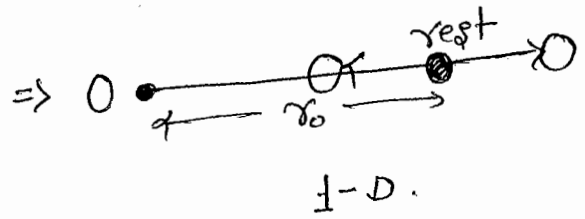
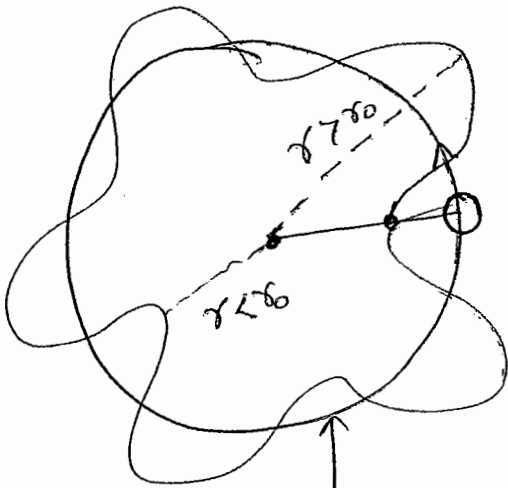
$$\text{K.E.} + \cancel{V(r)} = \frac{J^2}{2mr^2} + \cancel{V(r)}$$

$$\boxed{\text{K.E.} = \frac{J^2}{2mr^2}}$$

Always use for circular motion.

* Oscillation about stable State :-

A stable orbit is circular orbit because ($f_{eff} = 0$)



2-D. stable circular orbit

frequency of Oscillation :-

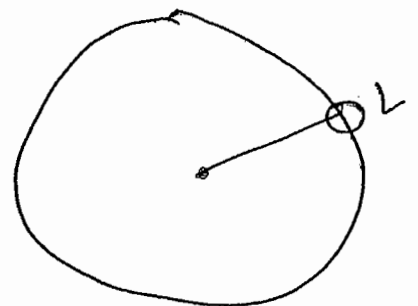
$$\omega = \sqrt{\frac{\text{force Constant}}{m}}$$

When $\omega =$ Radial freq. of angular oscillation.

$$\text{force Constant} = \left. \frac{d^2 V_{eff}}{dr^2} \right|_{r=r_0}$$

where $r_0 =$ radius of stable circular orbit.

B.A.-5
Q.32



Solⁿ

$$V(r) = \frac{-K}{r}$$

$$V_{\text{eff}} = V(r) + \frac{L^2}{2mr^2}$$

$$V_{\text{eff}} = \frac{-K}{r} + \frac{L^2}{2mr^2}$$

Let r_0 be radius of stable circular orbit -

$$\therefore f_{\text{eff}} = 0$$

$$\left. \frac{dV_{\text{eff}}}{dr} \right|_{r=r_0} = 0$$

$$\frac{K}{r_0^2} - \frac{L^2}{mr_0^3} = 0$$

$$\boxed{r_0 = \frac{L^2}{mk}}$$

$$\text{Hence force constant} = \left. \frac{d^2 V_{\text{eff}}}{dr^2} \right|_{r=r_0}$$

$$= \frac{-2K}{r_0^3} + \frac{3L^2}{mr_0^4}$$

$$= \frac{1}{r_0^3} \left(-2K + \frac{3L^2}{mr_0} \right)$$

$$= \frac{(mk)^3}{L^6} \left(-2K + \frac{3L^2 \times mk}{mL^2} \right)$$

$$= \frac{m^3 k^3}{L^6} (-2K + 3K) = \frac{m^3 k^4}{L^6}$$

Hence angular frequency :-

$$\omega = \sqrt{\frac{\text{force constant}}{m}}$$

$$= \sqrt{\frac{m^2 k^4}{L^6 m}} = \frac{m k^2}{L^3}$$

$$\boxed{\omega = \frac{m k^2}{L^3}} \quad \underline{\underline{\text{Ans}}}$$

Q. Graph of ' V_{eff} ' vs ' r ' :-

$$V_{\text{eff}} = V(r) + \frac{l^2}{2mr^2}$$

Consider following two limits

$r \rightarrow 0$, $r \rightarrow \infty$ and draw line in these two limits.

in most of the cases

if $r \rightarrow 0$ then higher power in denominator dominates
if $r \rightarrow \infty$ " lower " " " "

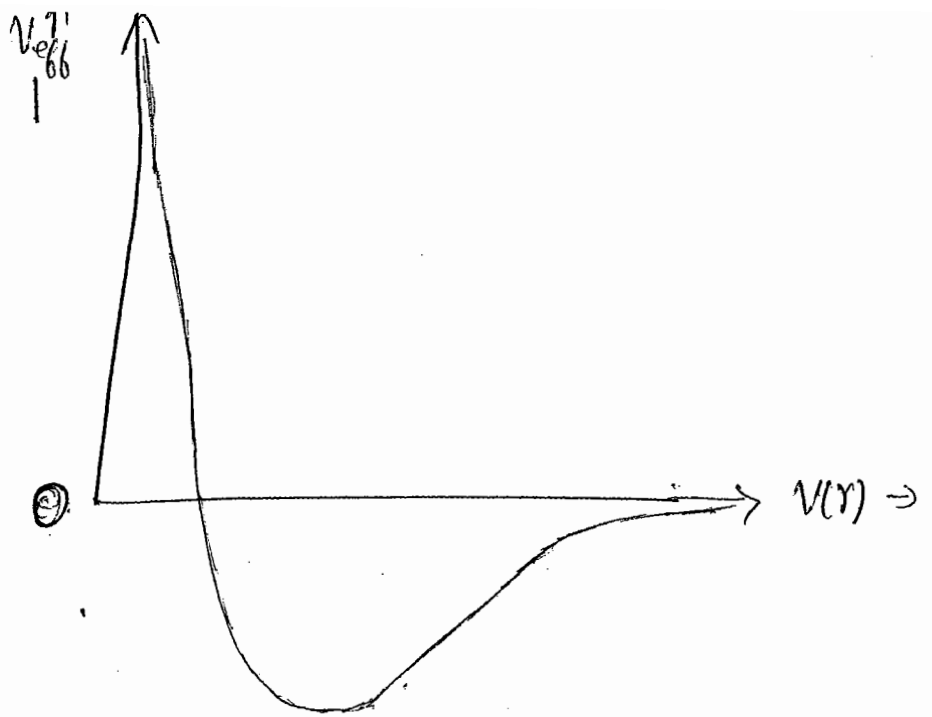
* if $r \rightarrow 0$ lowe power " " numerator " "
if $r \rightarrow \infty$ higher " " " "

Ans Plot V_{eff} for $V(r) = -\frac{K}{r}$

$$V_{\text{eff}} = -\frac{K}{r} + \frac{l^2}{2mr^2}$$

$$r \rightarrow 0, V_{\text{eff}} \rightarrow \infty$$

$$r \rightarrow \infty, V_{\text{eff}} \rightarrow 0 \quad (\text{from -ve side})$$



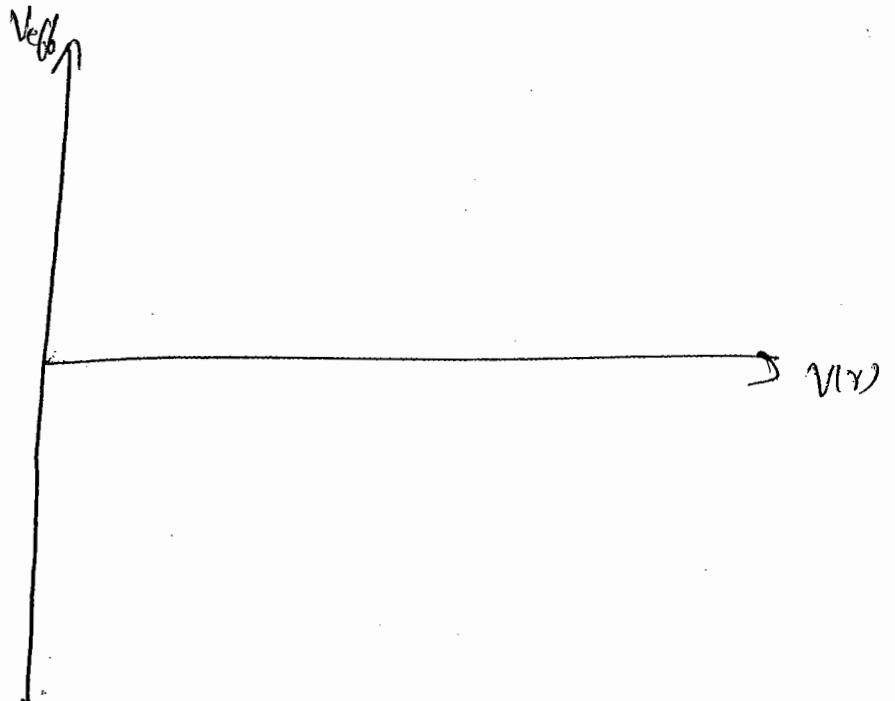
Q: V_{eff} for $V(r) = -\frac{K}{r^3}$

Soln: $V_{eff} = V(r) + \frac{L^2}{2mr^2}$

$$V_{eff} = -\frac{K}{r^3} + \frac{L^2}{2mr^2}$$

$$r \rightarrow 0, V_{eff} \rightarrow -\infty$$

$$r \rightarrow \infty, V_{eff} \rightarrow 0 \text{ (from +ve side)}$$

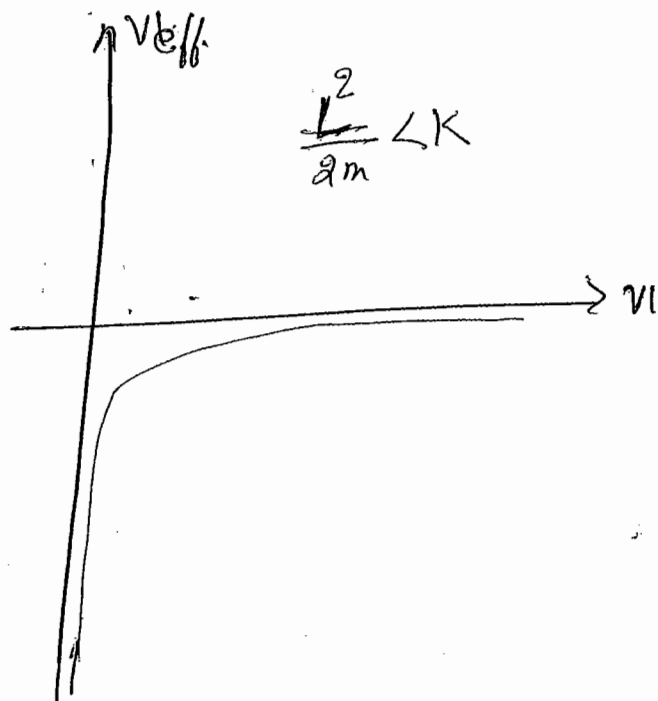
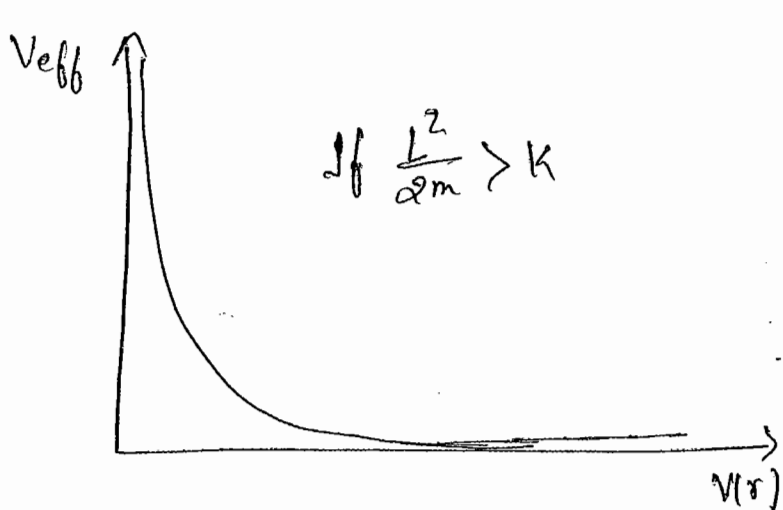


Q $V(r) = -\frac{K}{r^2}$

$$V_{eff} = -\frac{K}{r^2} + \frac{L^2}{2mr^2}$$

$$= \frac{1}{r^2} \left[\frac{L^2}{2m} - K \right]$$

when $\frac{L^2}{2m} > K$, when $\frac{L^2}{2m} < K$



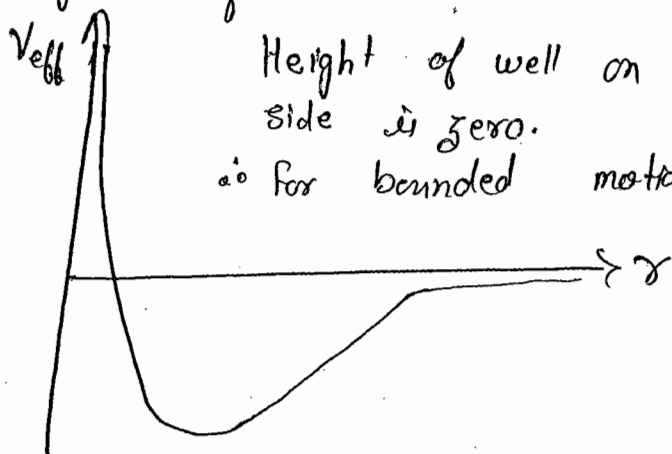
Bounded Motion

If there is a potential well in V_{eff} graph and energy of particle lying inside the well is less than minimum height of well then motion is bounded.

Example :-

$$V(r) = -\frac{K}{r}$$

$$V_{eff} = -\frac{K}{r} + \frac{L^2}{2mr^2}$$



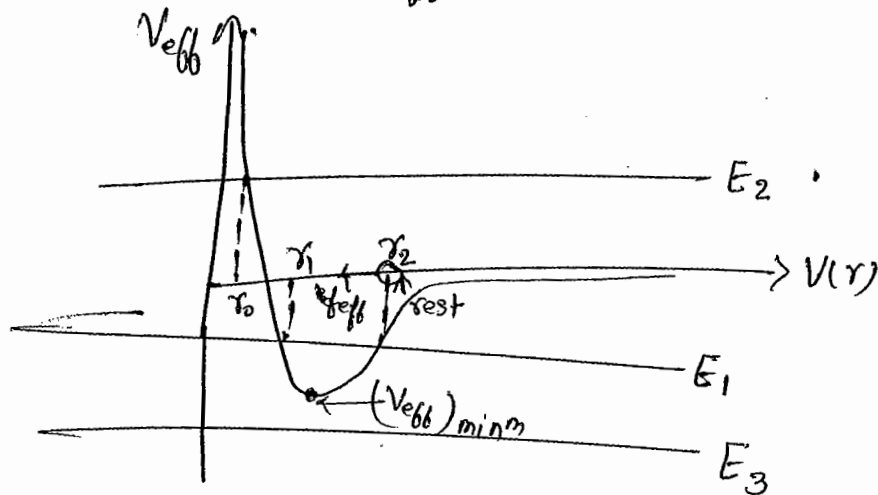
Height of well on right side is zero.
 ∴ For bounded motion $E < \dots$

$$E = \frac{1}{2} m \dot{r}^2 + V_{\text{eff}}$$

$$E - V_{\text{eff}} = \frac{1}{2} m \dot{r}^2 \leftarrow +ve \text{ or zero.}$$

$$E - V_{\text{eff}} \geq 0$$

$E \geq V_{\text{eff}} \Rightarrow$ Energy line will be above V_{eff} line.



- E_3 is not allowed.
- for energy E_1 particle moves b/w r_1 and r_2 (these are called turning point).
- for $E = E_2$ particle moves between r_0 and ∞

* Minimum allowed Energy :-

$$E_{\text{min}} = (V_{\text{eff}})_{\text{min}}$$

* Condition for stable orbit :-

$$\left. \frac{\partial V_{\text{eff}}}{\partial r} \right|_{r=r_0} = 0 \quad \text{--- (i)}$$

$$\left. \frac{\partial^2 V_{\text{eff}}}{\partial r^2} \right|_{r=r_0} > 0 \quad \text{--- (ii)}$$

In General :-

$$\text{If } F = kr^n$$

$$V(r) = - \int f(r) dr = \frac{-kr^{n+1}}{n+1}$$

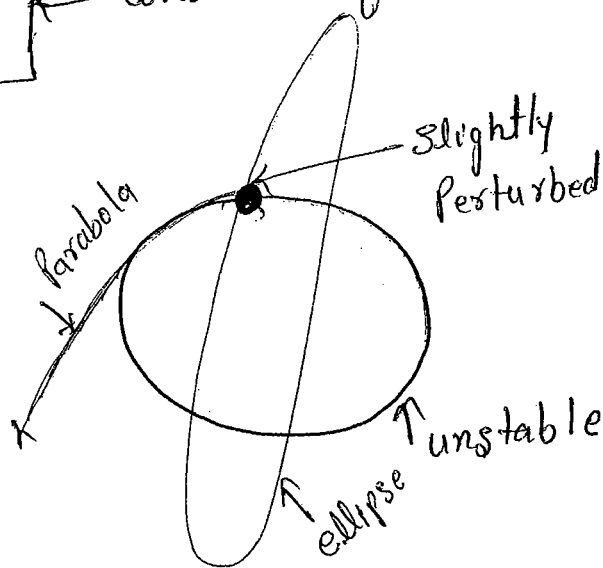
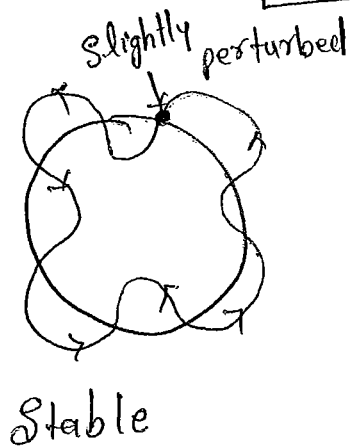
$$V_{\text{eff}} = V(r) + \frac{l^2}{2mr^2}$$

$$V_{\text{eff}} = \frac{-kr^{n+1}}{n+1} + \frac{l^2}{2mr^2}$$

from the above conditions ① & ② for above orbit, we get -

$$n > -3$$

Condition for stable orbit



* Condition for closed Orbit :-

After perturbation slightly

The object retrace it's path after some turns.
(Ultimately path should closed.)

If $F = kr^n$ then closed orbit is possible only for $n=1$ and $n=-2$.

[Forces should be attractive]

① Hook's Law

② Gravitational law. }

Q.5

Solⁿ

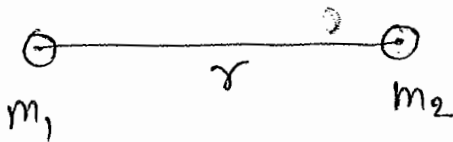
$$f = \frac{k}{r^{3-n}} = k r^{n-3}$$

$$n-3=1 \Rightarrow n=4$$

$$n-3=-2 \Rightarrow n=1$$

$$\boxed{n=4, 1} \quad \underline{\underline{Ans}}$$

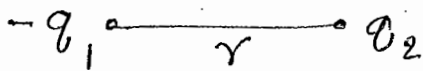
(*) Polar Equation of orbit under $V(r) = -\frac{k}{r}$



$$V(r) = \frac{-G m_1 m_2}{r} = -\frac{k}{r} \quad k = G m_1 m_2$$

$$f = \frac{-G m_1 m_2}{r^2}$$

$$\boxed{V(r) = -\int f(r) dr} = -\frac{k}{r}$$



$$V(r) = \frac{-q_1 q_2}{4\pi \epsilon_0 r} = -\frac{k}{r}$$

$$\Rightarrow \boxed{k = \frac{q_1 q_2}{4\pi \epsilon_0}}$$

Differential Equation of orbit:-

$$\frac{d^2 u}{d\theta^2} + u = \frac{-mf}{l^2 u^2}$$

$$\boxed{\frac{d^2}{d\theta^2} \left(\frac{1}{r}\right) + \frac{1}{r} = + \frac{mk}{l^2}}$$

$$f(r) = -\frac{\partial V}{\partial r}$$

$$= -\frac{k}{r^2}$$

$$= -k u^2$$

Solⁿ of this eqⁿ gives -

$$\boxed{r = \frac{l^2/mk}{1 \pm \left(\sqrt{1 + \frac{2El^2}{mk^2}}\right) \cos\theta}$$

→ Polar equation of orbit

* Polar form of conic section:-

$$r = \frac{A}{1 \pm e \cos\theta}$$

Compare to get -

$$\boxed{e = \sqrt{1 + \frac{2El^2}{mk^2}} = \text{eccentricity}}$$

e depends on E (energy)

$E > 0$, $e > 1$, orbit is hyperbola
 $-\frac{mk^2}{2l^2} < E < 0$, $e < 1$, orbit is ellipse.
 $E = 0$, $e = 1$, orbit is parabola
 $E = -\frac{mk^2}{2l^2}$, $e = 0$, orbit is circle.

} only for $V(r) = -\frac{k}{r}$

Q. A particle of mass m is projected from a height $h=R$ with speed $v = \sqrt{\frac{GM}{R}}$ where M is mass of earth, m is mass of particle. What is the path of the particle? in some direction?

Solⁿ "Note: If a particle is thrown in radial direction it will always be straight line."

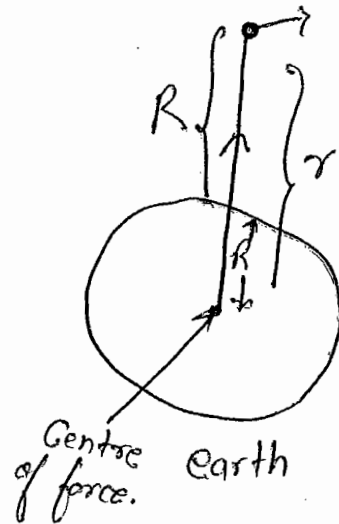
$$E = \frac{1}{2} mv^2 + V(r)$$

$$= \frac{1}{2} m \frac{GM}{R} - \frac{K}{r}$$

$$= \frac{GMm}{2R} - \frac{GMm}{2R}$$

$$\therefore \boxed{E = 0}$$

So path is parabola.

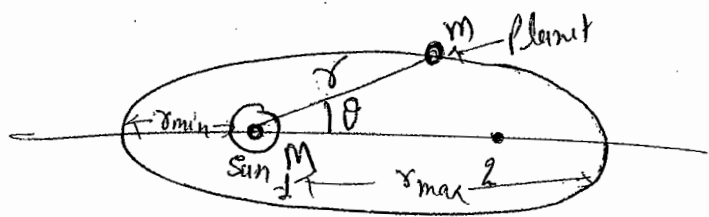


* Polar Equation of orbit :-

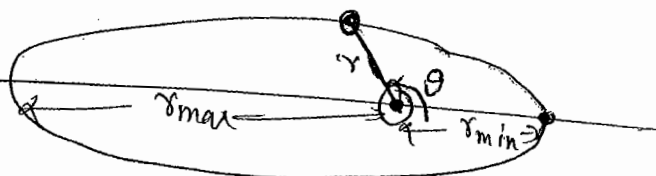
$$\boxed{r = \frac{l^2/mk}{1 \pm e \cos \theta}}$$

Elliptical Orbit :-

$$\boxed{r = \frac{l^2/mk}{1 - e \cos \theta}} \quad \text{--- (I)}$$



$$\boxed{r = \frac{l^2/mk}{1 + e \cos \theta}} \quad \text{--- (II)}$$



$$\left. \begin{array}{l} r_{\max} \text{ when } \theta = 0 \\ r_{\min} \text{ when } \theta = \pi \end{array} \right\} \text{ in eqn (i)}$$

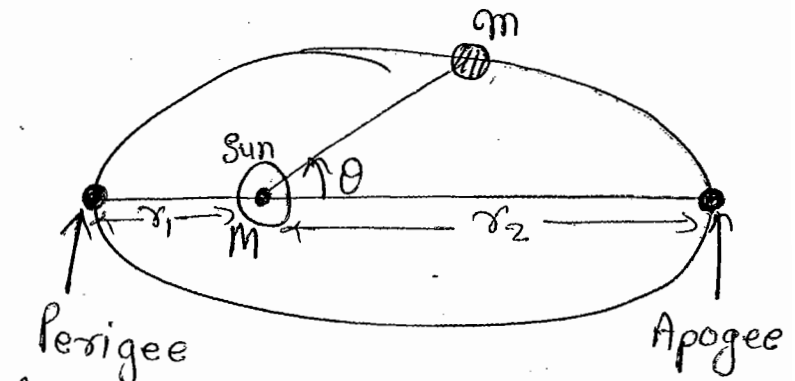
$$\left. \begin{array}{l} r_{\max} \text{ when } \theta = \pi \\ r_{\min} \text{ when } \theta = 0 \end{array} \right\} \text{ in eqn (ii)}$$

* Energy and angular momentum of planet:-
Say r_1 and r_2 are given

$$V(r) = -\frac{G M m}{r}$$

Polar equation of orbit-

$$r = \frac{l^2 / mk}{1 - \left(\sqrt{1 + \frac{2 E l^2}{m k^2}} \right) \cos \theta} \quad \text{(Perihelion)}$$



$$r_1 = \frac{l^2 / mk}{1 + \sqrt{1 + \frac{2 E l^2}{m k^2}}} \quad \text{--- (i)}$$

$$r_2 = \frac{l^2 / mk}{1 - \sqrt{1 + \frac{2 E l^2}{m k^2}}} \quad \text{--- (ii)}$$

$$1 + \sqrt{1 + \frac{2 E l^2}{m k^2}} = \frac{l^2}{mk} \cdot \frac{1}{r_1} \quad \text{--- (iii)}$$

$$1 - \sqrt{1 + \frac{2 E l^2}{m k^2}} = \frac{l^2}{mk} \cdot \frac{1}{r_2} \quad \text{--- (iv)}$$

$$\text{(iii)} + \text{(iv)} \quad 2 = \frac{l^2}{mk} \left(\frac{r_1 + r_2}{r_1 r_2} \right)$$

$$l = \sqrt{\frac{2r_1 r_2}{r_1 + r_2}} mk$$

$$l = m \sqrt{\frac{2GM r_1 r_2}{r_1 + r_2}}$$

Put l in eqⁿ (1) to get Energy -

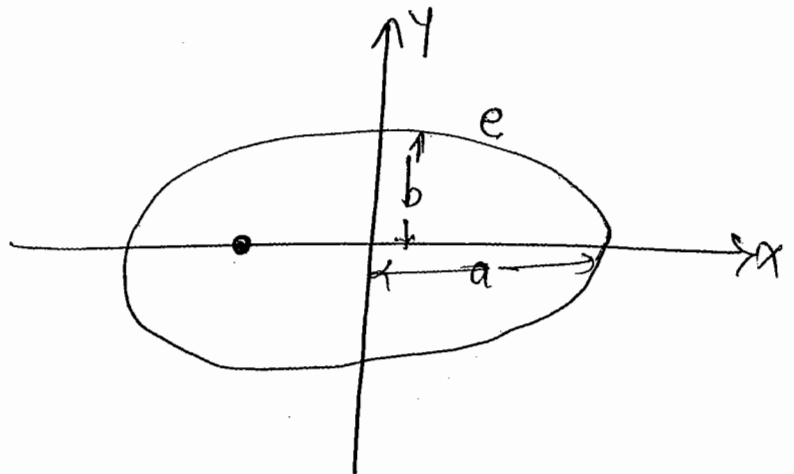
$$E = \frac{-GMm}{r_1 + r_2}$$

In terms of ellipse parameter :-

$$r_1 = (1-e)a$$

$$r_2 = (1+e)a$$

$$e = \sqrt{1 - \frac{b^2}{a^2}}$$



$$r_1 + r_2 = 2a$$

$$E = \frac{-GMm}{2a}$$

$$r_1 r_2 = (1-e^2)a^2$$

$$= \frac{b^2}{a^2} \cdot a^2$$

$$r_1 r_2 = b^2$$

$$L = m \sqrt{\frac{2GMb^2}{a}}$$

$$L = m \sqrt{\frac{GMb^2}{a}}$$

dependent on a and b .

* Maximum and Minimum Speed in elliptical orbit:-

$$L = m v_{\max} r_{\min} = m v_{\max} r_1$$

$$L = m v_{\min} r_{\max} = m v_{\min} r_2$$

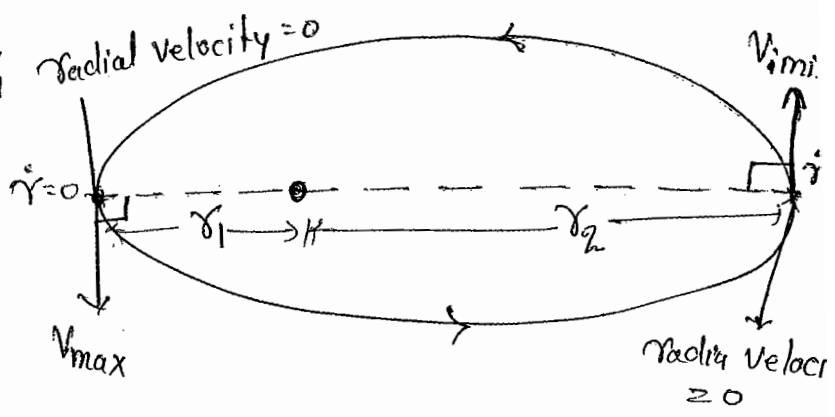
$$L = m \sqrt{\frac{2GM r_1 r_2}{r_1 + r_2}}$$

$$v_{\min} = \sqrt{\frac{GM r_1}{r_2 (r_1 + r_2)}}$$

$$v_{\max} = \sqrt{\frac{2GM r_2}{r_1 (r_1 + r_2)}}$$

$$r_1 = (1-e)a$$

$$r_2 = (1+e)a$$



$$v_{\min} = \sqrt{\frac{GM}{a} \left(\frac{1-e}{1+e} \right)}$$

$$v_{\max} = \sqrt{\frac{GM}{a} \left(\frac{1+e}{1-e} \right)}$$

Q. Planet is moving around the sun if ratio of maximum to minimum speed is 2. what is eccentricity of orbit?

Solⁿ

$$\frac{v_{\max}}{v_{\min}} = \frac{\sqrt{\frac{GM}{a} \left(\frac{1+e}{1-e} \right)}}{\sqrt{\frac{GM}{a} \left(\frac{1-e}{1+e} \right)}} = 2$$

$$\frac{1+e}{1-e} = 2$$

$$1+e = 2 - 2e$$

$$3e = 2 - 1$$

$$e = \frac{1}{3} \quad \text{Ans}$$

Q. 27

Solⁿ

Polar eqⁿ of orbit

→ Parabolic path
⇒ $E=0$, $e=1$
∴ highly elliptical
So we let it is
parabola

$$r = \frac{l^2/mk}{1+e\cos\theta}$$

1st case:

$$R_0 = \frac{l^2/mk}{1+1}$$

$$\left. \begin{array}{l} \cos\theta = \max^m \\ e=0, e=1 \end{array} \right\}$$

$$R_0 = \frac{l^2}{2mk} \Rightarrow \frac{l^2}{mk} = 2R_0$$

1st case:

Circular orbit :-

$$\therefore R_0 = \frac{l^2}{2mk} \Rightarrow \frac{l^2}{mk} = 2R_0$$

$$\therefore \frac{V_{\max}}{V_{\min}} = \frac{1+e}{1-e}$$

$$2 = \frac{1+e}{1-e}$$

Second case:- Circular orbit

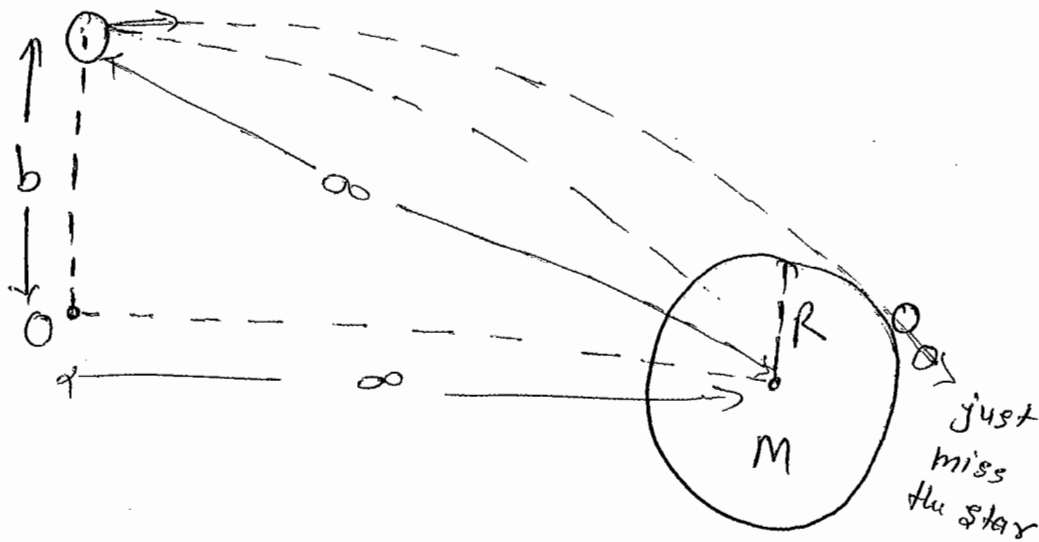
$$e=0$$

$$R_f = \frac{l^2/mk}{1+0}$$

$$\boxed{R_f = 2R_0} \quad R_m$$

Q.17

Solⁿ



$$r_{\min} = R$$

$$r = \frac{J^2/mk}{1 + e \cos \theta}$$

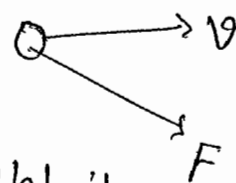
$$r_{\min} = \frac{J^2/mk}{1 + e}$$

$$R = \frac{J^2/mk}{1 + e}$$

$$R = \frac{J^2/mk}{1 + \sqrt{1 + \frac{2EL^2}{mK^2}}}$$

$$E = \frac{1}{2} m v^2 + V(r)$$

$$= \frac{1}{2} m v^2 - \frac{GMm}{\infty} = \frac{1}{2} m v^2$$



Velocity never cut the direction of force.

$$L = m v b$$

$$K = G M m$$

* Virial Theorem :-

It relates average value of Potential Energy and Kinetic Energy.

$$f = Kr^n$$

$$\boxed{\langle K.E. \rangle = \frac{n+1}{2} \langle P.E. \rangle} \leftarrow \text{Virial Theorem}$$

Ex -

for Gravitation :- $n = -2$

$$\boxed{\langle K.E. \rangle = -\frac{1}{2} \langle P.E. \rangle}$$

for Harmonic Oscillator :-

$$f = -Kr^1$$

$$n = 1$$

$$\boxed{\langle K.E. \rangle = \langle P.E. \rangle}$$

BA-5

Q.7

Potential energy $V = Kr^n$ then Relation b/w K.E. & P.E.

(a) $\langle T \rangle = \langle V \rangle$ (b) $\langle T \rangle = \frac{n}{2} \langle V \rangle$ (c) $\langle T \rangle = \frac{3}{2} \langle V \rangle$

(d) $\langle T \rangle = 2 \langle V \rangle$

Solⁿ

$$V = Kr^n$$

$$f = -\frac{\partial V}{\partial r} = -knr^{(n-1)}$$